WHY DO WE NEED THE NH MODEL ?

Acoustic-gravity waves

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INTRODUCTION

- The goal of this lecture is to illustrate the differences in wave propagation for the hydrostatic and nonhydrostatic systems.
- This has an impact on the response of models for orographic flows, as shown later by Jan Masek.
- Wave analysis also important as a basis for many numerical developments in models (SI scheme, stability studies...)
- In handbooks, the topic of waves is treated in z coordinates.
- Here, we decribe wave in mass-coordinate, to stick to model formulation

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Outline

1 What are waves?

- **2** Linearized system in σ coordinate
 - 3 Analysis of waves
- A Short discussion

- wave : oscillating perturbation around a stable equilibrium state in a medium
- linear wave : small oscillating perturbation ...
- The character and shape of linear waves depends on the choice of the stable equilibrium
- Complicated state \rightarrow dispersive propagation, non-uniform geometry etc...
- $\bullet~Simple~state \rightarrow more~regular~propagation~and~geometry$
- Hand-analysis of linear waves tractable only very simple equilibrium states.

We focus on small-scale atmospheric waves (neglect large-scale features as rotation f)

We focus on the simplest waves, i.e. on the simplest equilibrium state :

- resting (and stable \Rightarrow hydrostatic)
- isothermal
- dry, nonrotating, ...

Classically, to further simplify, we will assume an unbounded fluid (boundaries impose further constraints on waves) in 2D medium (x, σ) And of course we assume linear waves

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EE system in σ coordinate and model variables (d, \mathcal{P})

$$\frac{d\mathbf{V}}{dt} = -RT\nabla q - \frac{RT}{(1+\mathcal{P})}\nabla\mathcal{P} - \left(1+\mathcal{P}+\sigma\frac{\partial\mathcal{P}}{\partial\sigma}\right)\nabla\phi$$

$$\frac{d\mathbf{d}}{dt} = -\frac{g^2(1+\mathcal{P})}{RT}\left(\sigma\frac{\partial}{\partial\sigma}\right)\left(1+\sigma\frac{\partial}{\partial\sigma}\right)\mathcal{P}$$

$$+ d(\nabla\cdot\mathbf{V}-D_3) + \frac{g(1+\mathcal{P})}{RT}\left[\nabla w \cdot \left(\sigma\frac{\partial\mathbf{V}}{\partial\sigma}\right)\right]$$

$$\frac{dT}{dt} = -\frac{RT}{C_v} D_3 \frac{dP}{dt} = -(1+P) \left(\frac{C_p}{C_v} D_3 + \frac{\dot{\pi}}{\pi}\right) \frac{\partial q}{\partial t} = -\int_0^1 (\nabla \cdot \mathbf{V} + \mathbf{V} \cdot \nabla q) d\sigma'$$

Unbounded EE system in σ coordinate

Warning : all symbols in red imply vertical integrals with references to the surface.

Facultative part (sleeping allowed) :

Procedure to show that this system may apply to a vertically unbounded fluid :

- notice the surface $\phi = \phi_s$ is not necessarily a material one
- if not material, $\phi = \phi_s$ is just a reference, immaterial surface
- \bullet Then for unbounded medium, $\sigma \in [0,\infty],$ and $\sigma = 1$ at reference surface
- to remove integrals, simply differentiate vertically

Possible discussion tonight for those interested !

Unbounded EE system in σ coordinate

We use the short-hand notation $\partial_{\!\!\sigma}=\sigma \frac{\partial}{\partial\sigma}$

$$\phi = R \int_{\sigma}^{1} \frac{T}{1 + \mathcal{P}} \frac{d\sigma}{\sigma} \quad \Rightarrow \quad \partial_{\sigma} \phi = -\frac{RT}{1 + \mathcal{P}}$$

$$\begin{split} \frac{\dot{\pi}}{\pi} &= \mathbf{V} \cdot \boldsymbol{\nabla} q - \frac{1}{\sigma} \int_{0}^{\sigma} \left(\boldsymbol{\nabla} \cdot \mathbf{V} + \mathbf{V} \cdot \boldsymbol{\nabla} q \right) d\sigma \\ \Rightarrow & (1 + \partial_{\sigma}) \frac{\dot{\pi}}{\pi} = \partial_{\sigma} \mathbf{V} \cdot \boldsymbol{\nabla} q - \boldsymbol{\nabla} \cdot \mathbf{V} \end{split}$$

Unbounded EE system in σ coordinate

Equations for \boldsymbol{V} and $\boldsymbol{\mathcal{P}}$ become :

$$\partial_{\sigma} \frac{d\mathbf{V}}{dt} = -\partial_{\sigma} \left[RT \nabla q + \frac{RT}{(1+\mathcal{P})} \nabla \mathcal{P} \right] \\ + \left[1 + (1+\partial_{\sigma})\mathcal{P} \right] \nabla \frac{RT}{(1+\mathcal{P})} - \left[\partial_{\sigma} (1+\partial_{\sigma})\mathcal{P} \right] \nabla \phi \\ + \partial_{\sigma} \right) \frac{d\mathcal{P}}{dt} = -\frac{C_{\rho}}{C_{\nu}} (1+\partial_{\sigma}) \left[(1+\mathcal{P})D_{3} \right] \\ + \left[\nabla \mathbf{V} - \partial_{\sigma} \mathbf{V} \cdot \nabla q \right] - (1+\partial_{\sigma}) \left(\mathcal{P} \frac{\dot{\pi}}{\pi} \right)$$

Now we linearize...

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Linearization of the system

Basic state (resting, isothermal, homegeneous...)

$$\partial_{\sigma} \frac{d\mathbf{V}}{dt} = -\partial_{\sigma} \left[RT \nabla q + \frac{RT}{(1+\mathcal{P})} \nabla \mathcal{P} \right] \\ + \left[1 + (1+\partial_{\sigma})\mathcal{P} \right] \nabla \frac{RT}{(1+\mathcal{P})} - \left[\partial_{\sigma} (1+\partial_{\sigma})\mathcal{P} \right] \nabla \phi \\ + \partial_{\sigma} \right) \frac{d\mathcal{P}}{dt} = -\frac{C_{P}}{C_{V}} (1+\partial_{\sigma}) \left[(1+\mathcal{P})D_{3} \right] \\ + \left[\nabla \mathbf{V} - \partial_{\sigma} \mathbf{V} \cdot \nabla q \right] - (1+\partial_{\sigma}) \left(\mathcal{P} \frac{\dot{\pi}}{\pi} \right)$$

Terms in red are non-linear (neglected).

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Linearization

$$D_3 = \nabla \cdot \mathbf{V} + \mathsf{d} + \frac{(1+\mathcal{P})}{RT} \nabla \phi \cdot \partial_{\sigma} \mathbf{V} \implies D_3 \to D + \mathsf{d}$$

$$(d/dt) = (\partial/\partial t) + Advection \implies (d/dt) \rightarrow (\partial/\partial t)$$

$$\partial_{\sigma} \frac{\partial D}{\partial t} = -RT^* \partial_{\sigma} \nabla^2 \mathcal{P} + R\nabla^2 T - RT^* \nabla^2 \mathcal{P}$$
$$= R\nabla^2 T - RT^* (1 + \partial_{\sigma}) \nabla^2 \mathcal{P}$$
$$1 + \partial_{\sigma}) \frac{\partial \mathcal{P}}{\partial t} = -\frac{C_p}{C_v} (1 + \partial_{\sigma}) [D + d] + D$$

Linearization (cont'd)

Finally the linearized unbounded system writes :

$$\partial_{\sigma} \frac{\partial D}{\partial t} = R\nabla^{2}T - RT^{*}(1+\partial_{\sigma})\nabla^{2}\mathcal{P}$$

$$\gamma \frac{\partial d}{\partial t} = -\frac{g^{2}}{RT^{*}}\partial_{\sigma}(1+\partial_{\sigma})\mathcal{P}$$

$$\frac{\partial T}{\partial t} = -\frac{RT^{*}}{C_{v}}(D+d)$$

$$1+\partial_{\sigma})\frac{\partial \mathcal{P}}{\partial t} = -\frac{C_{p}}{C_{v}}(1+\partial_{\sigma})[D+d]+D$$

where γ is the marker of hydrostatic approximation

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Analysis of waves

The linear system admit solutions of the form : $\psi(x, \sigma, t) = \hat{\psi}(\sigma) \exp i(kx + \omega t)$ Hence, with "hats" dropped :

$$i\omega\partial_{\sigma}D = -k^{2}RT + k^{2}RT^{*}(1+\partial_{\sigma})\mathcal{P}$$
$$i\gamma\omega d = -\frac{g^{2}}{RT^{*}}\partial_{\sigma}(1+\partial_{\sigma})\mathcal{P}$$
$$i\omega T = -\frac{RT^{*}}{C_{v}}(D+d)$$
$$\omega(1+\partial_{\sigma})\mathcal{P} = D - \frac{C_{p}}{C_{v}}(1+\partial_{\sigma})(D+d)$$

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Analysis of waves

Eliminating for D and d :

$$-\omega^{2}\partial_{\sigma}D = -k^{2}R\left[-\frac{RT^{*}}{C_{v}}(D+d)\right] + k^{2}RT^{*}\left[D - \frac{C_{p}}{C_{v}}(1+\partial_{\sigma})(D+d)\right]$$
$$-\gamma\omega^{2}d = -\frac{g^{2}}{RT^{*}}\partial_{\sigma}\left[D - \frac{C_{p}}{C_{v}}(1+\partial_{\sigma})(D+d)\right]$$

i.e.
$$-\partial_{\sigma} \left[\omega^{2} - k^{2}c^{2}\right] D = k^{2}RT^{*} \left[\frac{R}{C_{\nu}} - \frac{C_{\rho}}{C_{\nu}}(1 + \partial_{\sigma})\right] d$$
$$\left[\gamma\omega^{2} + \frac{g^{2}}{RT^{*}}\frac{C_{\rho}}{C_{\nu}}(1 + \partial_{\sigma})\partial_{\sigma}\right] d = \frac{g^{2}}{RT^{*}} \left[1 - \frac{C_{\rho}}{C_{\nu}}(1 + \partial_{\sigma})\right] \partial_{\sigma} D$$

where
$$c^2 = (C_p/C_v)RT^3$$

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Combining the two latter equations (after some eliminations) :

$$\left\{\gamma\omega^4 - \omega^2 c^2 \left[\gamma k^2 - \frac{(1+\partial_\sigma)\partial_\sigma}{H^2}\right] + k^2 N^2 c^2\right\} d = 0$$

where $N^2 = (g^2/C_p T^*)$, and $H = RT^*/g$

Solutions have the form : $d(\sigma) = d_0 \sigma^{(i\nu H - 1/2)}$ with :

$$\gamma \omega^4 - c^2 \left[\gamma k^2 + \nu^2 + 1/4H^2 \right] + k^2 N^2 c^2 = 0$$

Case $\gamma = 0$ (Hydrostatic equations) :

The 2 frequencies of solutions for a given (k, ν) geometry are :

$$\omega^2 = \frac{k^2 N^2}{\nu^2 + 1/4H^2}$$

These represent gravity waves

Case $\gamma = 1$ (Euler equations) :

The 4 frequencies of solutions for a given (k, ν) geometry are :

$$\omega^{2} = \frac{1}{2} \left[c^{2} \left(k^{2} + \nu^{2} + 1/4H^{2} \right) \pm \sqrt{c^{4} \left(k^{2} + \nu^{2} + 1/4H^{2} \right)^{2} - 4k^{2}N^{2}c^{2}} \right]$$

These represent "gravity-acoustic" and "acoustic-gravity" waves

Special case "gravity" for
$$(k \ll \nu c/N)$$
, e.g. $(k \approx \nu)$

$$\omega^2 = \frac{k^2 N^2}{k^2 + \nu^2 + 1/4H^2}$$

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Compare H and NH "gravity" waves for $k \ll \nu c/N$



For hydrostatic systems :

- Gravity waves with aspect ratio pprox 1 will be considerably distorted
- Orographic (stationary) gravity waves will not radiate energy in the right direction (\longrightarrow Jan's talk)