# **Turbulence closure** for stably stratified geophysical flows

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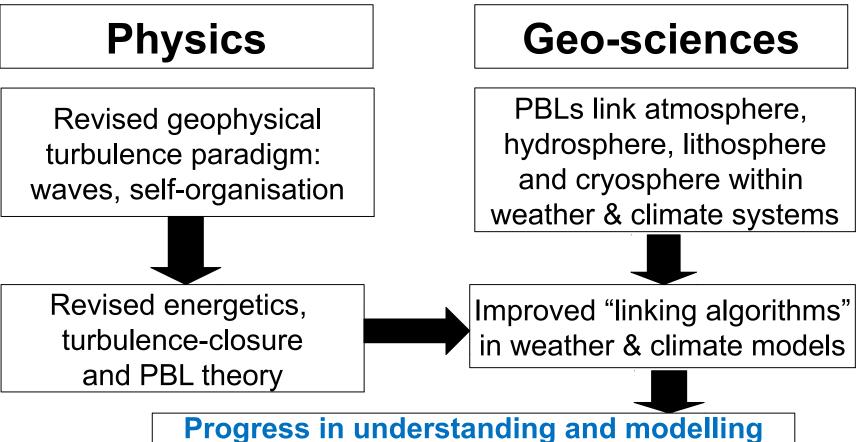
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# Geophysical turbulence and planetary boundary layers (PBLs)



weather & climate systems



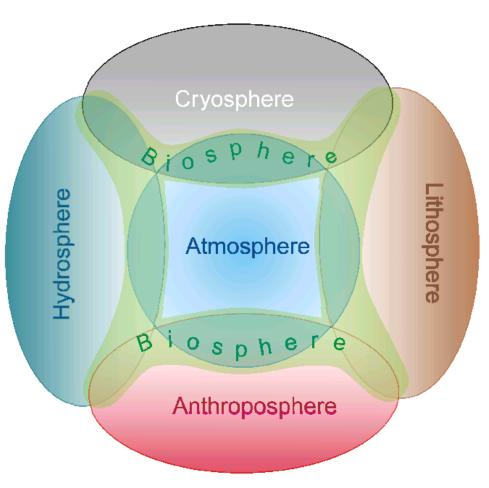


# **Geospheres in climate system**

Turbulence performs vertical transports of energy, matter and momentum in fluid geospheres

Atmosphere, hydrosphere, lithosphere and cryosphere are coupled through turbulent planetary boundary layers PBLs (dark green lenses)

PBLs include 90% biosphere and entire anthroposphere







#### Role of planetary boundary layers (PBLs): TRADITIONAL VIEW

Surface fluxes at interface between AIR

and

WATER (or LAND)

fully characterise interaction between **ATMOSPHERE and OCEAN/LAND** 

Monin-Obukhov similarity theory (1954) (conventional framework for determining surface fluxes in operational models) disregards non-local features of convective and long-lived stable PBLs



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# **Role of PBLs: MODERN VIEW**

Deep ocean **Oceanic PBI** (upper mixed layer) Atmospheric PBL phere

Because of very stable stratification in atmosphere and ocean beyond PBLs and convective zones, density increments inherent at PBL outer boundaries prevent entities delivered by surface fluxes (or emissions) to efficiently penetrate from PBL into free atmosphere or deep ocean.

Hence PBL heights and fluxes due to entrainment at PBL outer boundaries essentially control extreme weather events

(e.g., <u>heat waves</u> associated with convection; or strong stable stratification triggering <u>air pollution</u>).

This concept (equally relevant to hydrosphere) requires knowledge of <u>PBL height/depth</u> and <u>turbulent entrainment</u>

in numerical weather prediction, air/water quality and climate modelling.





# Very shallow boundary layer separated form the free atmosphere by capping inversion



PBL height visualised by smoke blanket (Johan The Ghost, Wikipedia) Capping inversion prevents PBL – free flow exchange



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#### Main stream in turbulence-closure theory

**Boussinesq** (1877) <u>Turbulent transfer is basically similar to molecular transfer</u> but much more efficient  $\rightarrow$  <u>down-gradient transport</u>  $\rightarrow$  *K*-theory  $\rightarrow$  eddy viscosity, conductivity, diffusivity

Richardson (1920, 1922) stratification (*Ri*), concept of forward energy cascade

Keller & Fridman (1924) a chain of budget equations for statistical moments <u>Problem</u>: to express the higher-order moments through the lower-order moments

**Prandtl** (1930s) mixing length  $l \sim z$ , velocity scale  $uT \sim ldU/dz$ , viscosity  $K \sim luT$ 

Kolmogorov (1941) quantified the cascade, closure as a problem of energetics:

- budget equation for <u>turbulent kinetic energy</u> (TKE)
- <u>TKE dissipation rate</u> expressed through the turbulent-dissipation <u>length scale</u>  $uT \sim (K\Im T)1/2, K \sim l \varepsilon uT$  underlies further developments through 20th century

**Obukhov** (1946) TKE-closure extended to stratified flows, **Obukhov length scale** *L* 

Monin & Obukhov (1954) alternative  $\rightarrow$  similarity theory for the surface layer z/L

**Mellor & Yamada** (1974) hierarchy of *K*-closures  $\rightarrow$  <u>turbulence cut-off problem</u>





#### **Turbulence cut-off problem**

Buoyancy  $b = (g/\rho 0)\rho$  (g – acceleration due to gravity,  $\rho$  –density)

Velocity shear S = dU/dz (U-velocity, z - height)  $Ri = \frac{db/dz}{(dt/dz)^2}$ 

**<u>Richardson number</u>** characterises static stability:

the higher Ri (or z/L), the stronger suppression of turbulence

#### Key question What happens with turbulence at large *Ri*?

<u>Traditional answer</u> Turbulence degenerates, and at Ri exceeding a critical value (Ricritical < 1) the flow inevitably becomes laminar (Richardson, 1920; Taylor, 1931; Prandtl, 1930,1942; Chandrasekhar, 1961;...)

In fact field, laboratory and numerical (LES, DNS) experiments show that <u>GEOPHYSICAL</u> (very high *Re*) turbulence is maintained up to  $Ri \sim 102$ Modellers were forced to <u>VIOLENTLY</u> preclude the turbulence cut-off





#### **Milestones**

Prandtl-1930's followed Boussinesq's idea of the down-gradient transfer (*K*-theory), determined  $K \sim luT$ , and expressed uT heuristically through the mixing length l

Kolmogorov-1942 (for neutrall stratication) followed Prandtl's concept of eddy viscosity  $KM \sim luT$ ; determined uT = (TKE)1/2 through TKE budget equation with dissipation  $\varepsilon \sim (TKE)/tT \sim (TKE)3/2/l\varepsilon$ ; and assumed  $l\varepsilon \sim l$  (grounded in neutral stratification)

Obukhov-1946 and then the entire turbulence community extended Kolmogorov's closure to <u>stratified flows</u> keeping it untouched, except for inclusion of the buoyancy term in the TKE equation. <u>Its sole use has caused cutting off TKE in supercritical stable stratification</u>

This approach, **missed** <u>turbulent potential energy</u> (TPE) and its interaction with TKE); overlooked inapplicability <u>of Prandtl's relation</u>  $K \sim luT$  to the eddy conductivity KH; and disregarded <u>principal deference between</u>  $l\varepsilon$  and l

For practical applications Mellor and Yamada (1974) developed <u>corrections preventing</u> <u>unacceptable turbulence cut-off</u> in "supercritical" static stability



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## Energy- & flux-budget (EFB) closure (2007-12)

**Budget equations for major statistical moments** 

Turbulent kinetic energy (TKE)EKTurbulent potential energy (TPE)EPVertical flux of temperature $Fz = \langle \theta w \rangle$  [or buoyancy (g/T)Fz]Vertical flux of momentum $\tau iz = \langle uiw \rangle$  (i = 1,2)

**Relaxation equation for the dissipation time scale**  $tT = EK/\varepsilon K = l(EK)-1/2$ 

Accounting for TPE  $\rightarrow$  vertical heat flux (that "killed" TKE in Kolmogorov type closures) <u>drops out from the equation for total turbulent energy</u> (TTE = TKE + TPE) <u>Heat-flux budget equation</u>  $\rightarrow$  imposes a limit on the vertical heat flux and assures <u>self-preservation of turbulence</u>  $\rightarrow$  <u>no</u> Ri-<u>critical in the energetic sense</u>

#### Disclosed two principally different regimes of stably stratified turbulence <u>"Strong turbulence" in boundary layer flows</u>

with  $KM \sim KH$  at Ri < Ric "Weak turbulence" in the free atmosphere with  $PrT = KM/KH \sim 4 Ri$  at Ri >>Ric

<u>MOS theory disregards</u> weak turbulence at z/L >>1 and yields artefact *Ric* <u>height</u> = the boundary between <u>strong</u>- and <u>weak-turbulence</u> regimes





PBL

# Turbulent potential energy – analogy to Lorenz (1955) available potential energy

Buoyancy fluctuation proportional to displacement of fluid particle

$$b' = \frac{g}{\rho_0} \rho' = \frac{g}{\rho_0} \frac{\partial \langle \rho \rangle}{\partial z} z' = N^2 z'$$

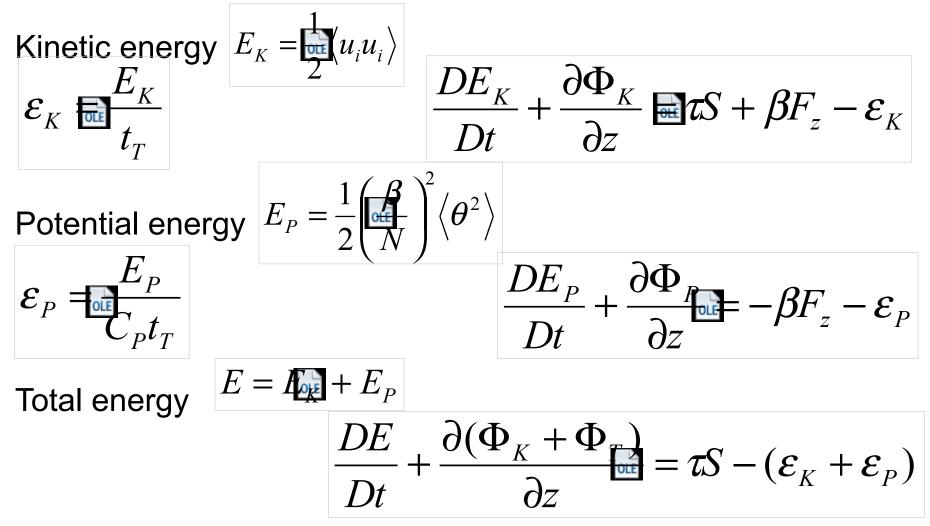
Potential energy (per unit mass) proportional to squared temperature

$$(E_{P})' = \frac{1}{z'} \int_{z}^{z+z'} b' z \, dz$$
  
=  $\frac{1}{2} \frac{(b')^{2}}{N^{2}} = \frac{1}{2} \left(\frac{\beta}{N}\right)^{2} (\theta')^{2} = \left(\frac{\beta}{N}\right)^{2} (E_{\theta})'$ 





#### **Turbulent energy budgets**



Buoyancy flux  $\beta Fz$  drops out from the turbulent total energy budget



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# Budget equation for the vertical<br/>turbulentturbulentflux of potential<br/> $F_z = 10$ temperature

$$\frac{DF_z}{Dt} + \frac{\partial}{\partial z} \Phi_z^{(F)} = C_\theta \beta \omega^2 \left\langle -2E_z \frac{\partial \Theta}{\partial z} - \frac{F_z}{C_F t_T} \right\rangle$$

The "pressure term" is shown to be proportional to the mean squared

fluctuation of potential temperature:

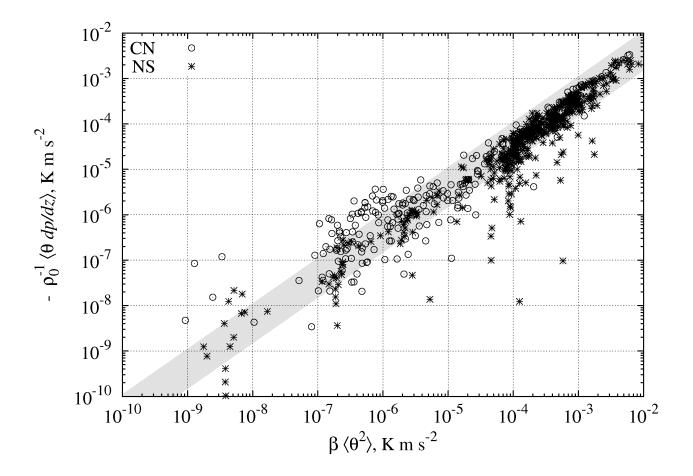
$$\frac{1}{\rho_0} \left\langle \theta \frac{\partial p}{\partial z} \right\rangle \sim \beta \left\langle \theta^2 \right\rangle$$

On the r.h.s. of the equation, 1st term (generation of positive heat flux) counteracts to 2nd term (generation of negative heat flux) and yields **self-preservation of turbulence in very stable stratification** 





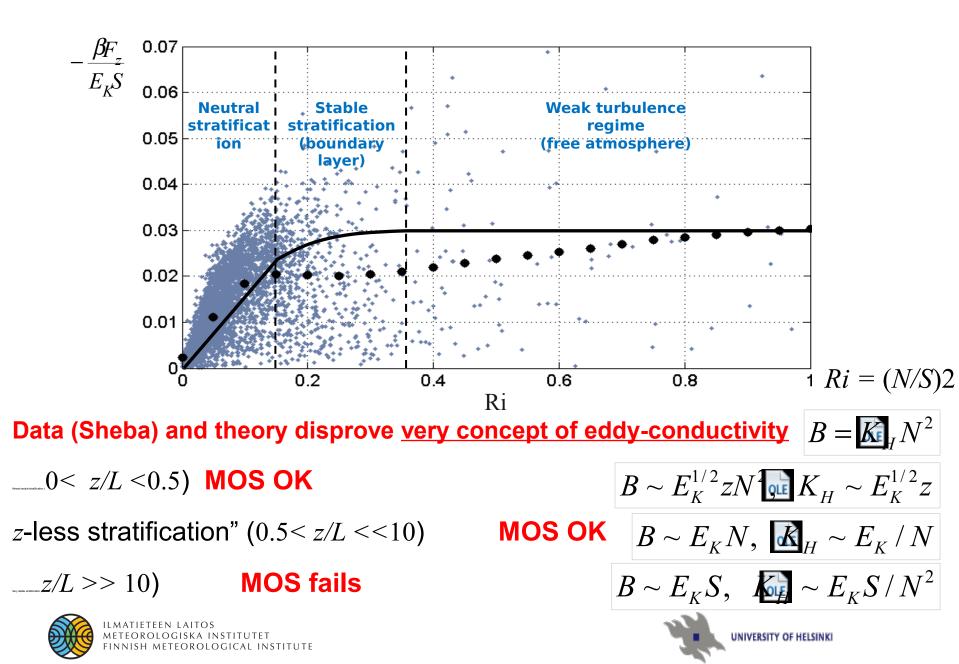
# **LES verification of our parameterization of the pressure term** $\rho_0^{-1} \langle \theta \partial p / \partial z \rangle \sim \beta \langle p^2 \rangle$







#### *Ri*-dependence of the buoyancy flux $B = \beta Fz$



## **Turbulent dissipation time and length scales**

By definition, time scale 
$$t_T \equiv \Box / \mathcal{E}_K$$
 and length scale  $l \equiv \Box / \mathcal{E}_T$   
The steady-state TKE budget  $\tau S + \beta F_z \equiv \tau S (1 - \Box i_f) = \mathcal{E}_K \equiv -\frac{E_K}{t_{TE}}$   
Flux Ri  
number  $Ri_f \equiv \frac{-\beta F_z}{\tau S} = \frac{\tau^{1/2}}{SL}$  Obukhov  $L = \frac{\tau^{3/2}}{-\beta F_z}$   $Ri_f - \Box R_\infty < 1$   
Shear: neutral  $S = \frac{\tau^{1/2}}{kz}$ , extreme stable (TKE)  $S \rightarrow \frac{-\beta F}{R_\infty \tau} = \frac{\tau^{1/2}}{R_\infty L}$   
Interpolation yields empirical  
law valid in any stratification  $S = \frac{\tau^{1/2}}{kz} \left(\Box + \frac{k}{R_\infty} \frac{z}{L}\right) \frac{k/R_\infty \tau}{k/R_\infty T} = \frac{\tau}{R_\infty L}$ 

law with the TKE equation yields

$$t_{TE} = \frac{kz}{E_K^{1/2} + C_\Omega \Omega z} \left(\frac{E_K}{\tau}\right)^{3/2} \frac{1 - Ri_f / R_\infty}{1 - Ri_f}$$

where kz plays the role of a "master length scale"



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#### **Relaxation equation for dissipation time scale**

Evolution of tT is controlled by

tendency towards equilibrium

$$t_T \rightarrow t_{TE}$$

counteracted by **distortion** due to non-stationary processes and heterogeneity causing mean-flow and turbulent transports.

This **counteraction** is described by **RELAXATION EQUATION** 

$$\frac{Dt_T}{Dt} - \frac{\partial}{\partial z} K_T \frac{\partial t_T}{\partial z} = -C_R \left(\frac{t_T}{t_{TE}} - 1\right)$$

 $C \square 1$  relaxation constant (differs for increasing/decreasing regimes) KT is the vertical turbulent exchange coefficient (~ to eddy viscosity)





## **EFB closure and M-O similarity theory (MOST)**

New physics behind known relation

Combined with the flux Richardson number

Yields **CONVERTOR** between *Rif* and z/L

EFB theory yields **CONVERTOR** between *Rif* and *Ri*:  $Ri_f = \prod_{P r_T(Ri)} Ri_{T}$ 

where 
$$Pr_T \approx \mathbb{R} = 0.8$$
 at  $Ri <<1$ ,  $Pr_T \approx Ri\mathbb{R}_{\infty} \approx 4Ri$  at  $Ri >>1$ 

See below empirical *Ri*-dependence of the turbulent Prandtl number





$$S = \frac{dU}{dz} = \frac{\tau_{kz}^{1/2}}{kz} \left( 1 + \frac{k}{R_{\infty}} \frac{z}{L} \right)$$
  
er 
$$Ri_{f} = \frac{\beta F_{z}}{\tau S}$$

$$Ri_f = \frac{kz/L}{1+kR_{\infty}^{-1}z/L}$$

$$Ri_f = \frac{k^2 / L}{1 + kR_{\infty}^{-1}z / L}$$

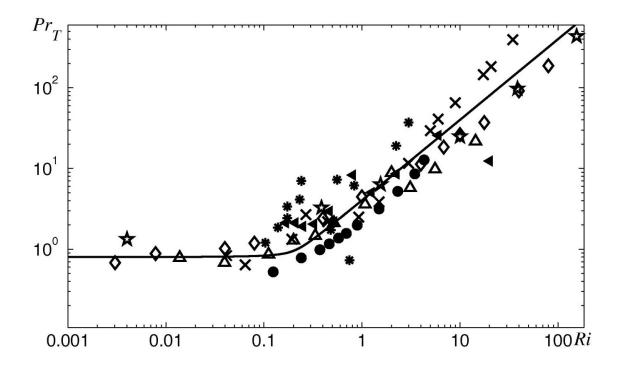
#### Major results

- The concept of *turbulent potential energy* (Z et al., 2007) analogous to Lorenz's *available potential energy* (both ~ squared density)
- New vision and relaxation equation for dissipation time scale
- Disproving <u>erroneous conclusion</u> that at high-*Re* the flow becomes laminar at *Ri* exceeding critical *Ric* ~ 0.25-1. In fact, it demarcates: reknown <u>Strong turbulence</u> with  $KM \sim KH$  at Ri < Ric typical of PBLs new <u>weak turbulence</u> with  $PrT = KM/KH \sim 4Ri$  at Ri >>Ric in free flow
- <u>Hierarchy of closure models</u> of different complexity for use in research and operational modelling
- Revision of Monin-Obukhov similarity theory
- Field, lab and numerical (LES, DNS) experiments confirm EFB theory for conditions typical of free atmosphere and deep ocean up to  $Ri \sim 103$





#### **Turbulent Prandtl number** PrT = KM/KH **versus** Ri



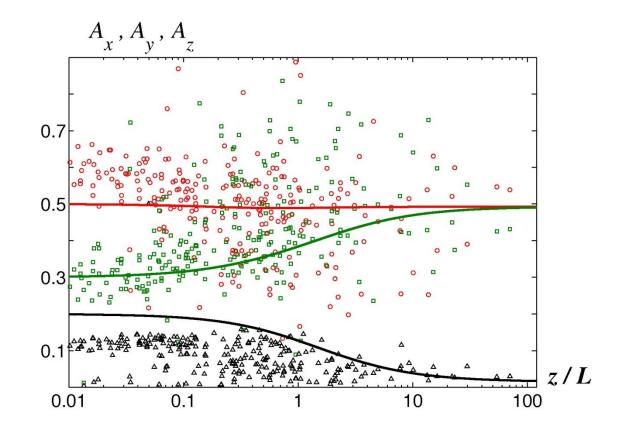
Atmospheric data:  $\blacktriangleleft$  (Kondo et al., 1978),  $\ast$  (Bertin et al., 1997); laboratory experiments:  $\land$  (Rehmann & Koseff, 2004),  $\diamondsuit$  (Ohya, 2001),  $\bullet$  (Strang & Fernando, 2001); DNS:  $\ast$  (Stretch et al., 2001); and LES:  $\triangle$  (Esau, 2009). The curve sows our EFB theory. The "strong" turbulence ( $PrT \approx 0.8$ ) and the "weak" turbulence ( $PrT \sim 4 Ri$ ) match at  $Ri \sim 0.25$ .

#### **MOST** assumes *PrT* = *constant*





#### **Longitudinal** Ax, transverse Ay & vertical Az TKE shares vs. z/L

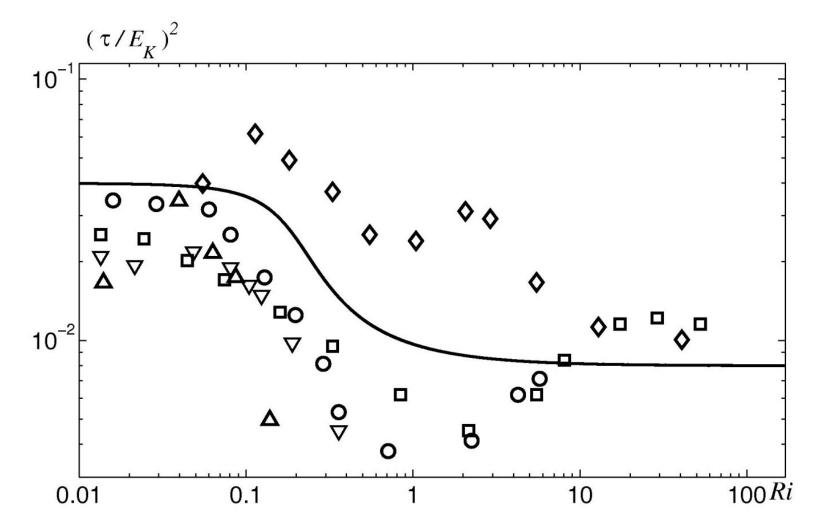


Experimental data from Kalmykian expedition 2007 of the Institute of Atmospheric Physics (Moscow). Theoretical curves are plotted after the EFB theory. The **traditional** "**return-to-isotropy**" model overlook the stability dependence of *Ay* clearly seen in the Figure. The strongest stability, z/L = 100, corresponds to Ri = 8.





#### Dimensionless vertical flux of momentum: two plateaus corresponding to the *strong* and *weak* turbulence regime MOST assumes $\tau/EK = constant$

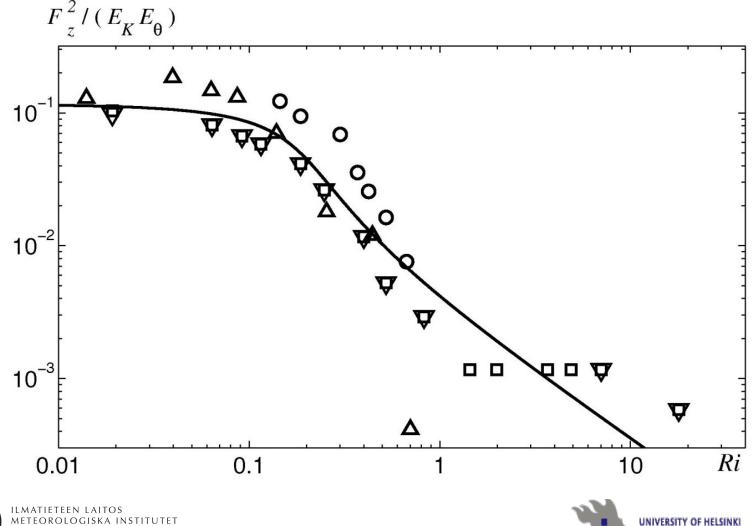






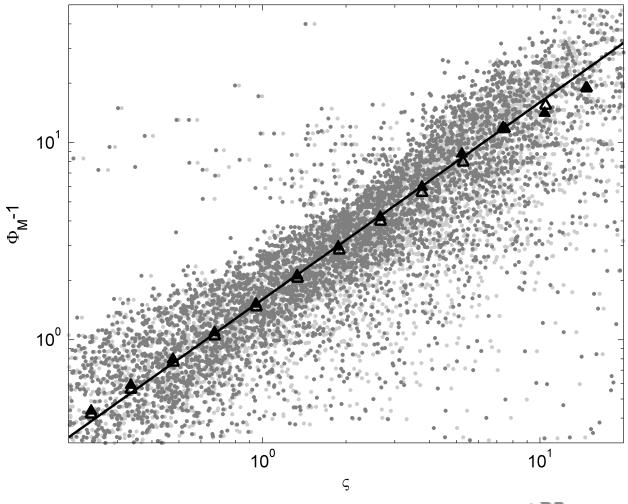
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#### Dimensionless heat flux: practically constant in *strong* turbulence and sharply decreases in *weak* turbulence **MOST assumes** $F_z/(EK E\theta)1/2 = constant$



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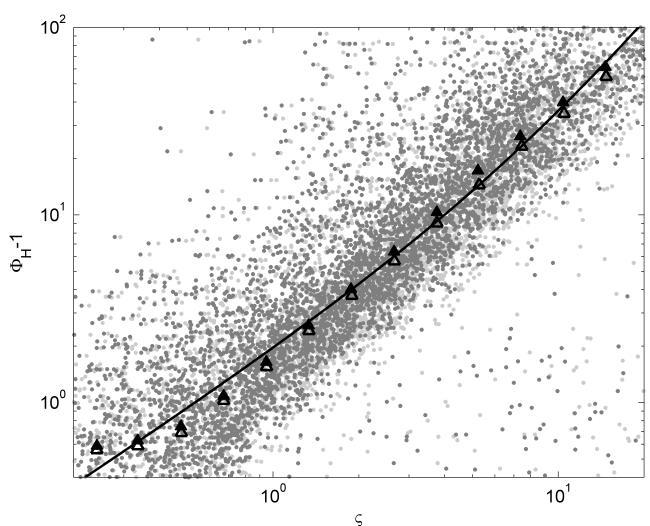
#### **Dimensionless velocity gradient** $\Phi_M = (kz / M) / (\partial U / \partial z)$ versus $\zeta = z/L$ after LES (dots) and the EFB model (curve) **MOST OK**



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Dimensionless temperature gradient versus  $\zeta = z/L$  after LES (dots) and the  $\mathbb{E}FB$  moder (curve)





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#### General closure model: energy&flux equations

$$\begin{array}{ll} \text{Kinetic energy} & \frac{DE_{K}}{Dt} - \frac{\partial}{\partial z} K_{E} \frac{\partial E_{K}}{\partial z} = \tau_{i3} \frac{\partial U_{i}}{\partial z} + \beta F_{z} - \frac{E_{K}}{t_{T}} \\ \text{Potential energy} & \frac{DE_{P}}{Dt} - \frac{\partial}{\partial z} K_{E} \frac{\partial E_{P}}{\partial z} = -\beta F_{z} - \frac{E_{P}}{C_{P}t_{T}} \\ \text{Momentum flux} & \frac{D\tau_{i3}}{Dt} - \frac{\partial}{\partial z} K_{FM} \frac{\partial \tau_{i2}}{\partial z} = -2E_{z} \frac{\partial U_{i}}{\partial z} - \frac{\tau_{i3}}{C_{z}t_{T}} \\ \text{\Theta-flux} & \frac{DF_{z}}{Dt} - \frac{\partial}{\partial z} \mathbf{M}_{FH} \frac{\partial F_{z}}{\partial z} = -2(E_{z} - C_{\theta} \mathbf{I}_{\Theta}) \frac{\partial \Theta}{\partial z} - \frac{F_{z}}{C_{F}t_{T}} \end{array}$$

Turbulent exchange coefficients for energies and fluxes are taken proportional to the eddy viscosity

$$K_E / C_E = K_{FM} / C_{FM} = K_H / C_{FH} = K_T / C_T = E_z t_T$$





#### **General closure model:**

## **Vertical TKE**

To characterise stability we use, instead of *Rif*, the energy-ratio

 $\Pi = \bigotimes_{K} / E_{K} \quad \{\text{in the steady-state } \Pi = CPRif / (1 - Rif) \} \text{ and} \\ \text{employ} \quad E_{z} \bigotimes_{K} E_{K} \quad E_{z} \underbrace{\Gamma}_{K} \tau$ 

our steady-state solution to express and as

universal functions of  $\Pi$  determined from our prognostic equations

## **Dissipation time scale**

Similarly, we express the equilibrium time scale TK through  $\Pi$  $t_{TE} = \frac{1}{E_K^{1/2} + C_\Omega \Omega z} \left(\frac{E_K}{\tau}\right)^2 \left(1 - \frac{1}{\Pi_\infty}\right)^2$ 

and determine  $\frac{Dt_T}{Dt} = \frac{1}{2} \frac{1}{2}$ 



## **Optimal closure model**

For operational modelling	we
recommend as optimal the model based on 3 prognost	ic equations
for:	- the two
turbulent energies $EK$ and $EP$	- and the
dissipation time scale $tT$	- in
combination with diagnostic eddy viscosity & eddy cond	uctivity
Advantages of the EFB closures:	
<ul> <li>consistent energetics with no Ri-critical</li> </ul>	-
advanced concept of the turbulent dissipation time scal	e -
"energy stratification parameter" preventing artificial ext	tremes -
essential anisotropy of turbulence	-

generally non-gradient and non-local turbulent transports





#### Conclusions

- EFB turbulence closure → new vision and modelling of geophysical stably stratified turbulence
- No Ric in the energetic sense: experimental data confirm this conclusion up to  $Ri \sim 103$
- Instead: *Ri* ~ 0.2-0.3 (hydrodynamic instability limit) separates regimes of "strong" and "weak" turbulence → <u>the boundary between PBL and free atmsophere</u> → another view at the PBL height
- MOS is applicable to the "strong" turbulence regime typical of boundary layer flows but inapplicable to "weak turbulence" typical of free atmosphere / ocean



